# BLOW-UP CRITERIA AND INSTABILITY OF STANDING WAVES FOR THE FRACTIONAL SCHRÖDINGER POISSON EQUATION 

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#### Abstract

In this article, we consider blow-up criteria and instability of standing waves for the fractional Schrödinger-Poisson equation. By using the localized virial estimates, we establish the blow-up criteria for non-radial solutions in both mass-critical and mass-supercritical cases. Based on these blow-up criteria and three variational characterizations of the ground state, we prove that the standing waves are strongly unstable. These obtained results extend the corresponding ones presented in the literature.


## 1. Introduction

In recent years, there has been a great deal of interest in using fractional Laplacians to model physical phenomena. By extending the Feynman path integral from the Brownian-like to the Lévy-like quantum mechanical paths, Laskin [23, 24] used the theory of functionals over functional measure generated by the Lévy stochastic process to introduce the fractional nonlinear Schrödinger equation (NLS)

$$
\begin{equation*}
i \partial_{t} \psi-(-\Delta)^{s} \psi+f(\psi)=0 \tag{1.1}
\end{equation*}
$$

where $i^{2}=-1,0<s<1$ and $f(\psi)$ is the nonlinearity. The fractional differential operator $(-\Delta)^{s}$ is defined by $(-\Delta)^{s} \psi=\mathcal{F}^{-1}\left[|\xi|^{2 s} \mathcal{F}(\psi)\right]$, where $\mathcal{F}$ and $\mathcal{F}^{-1}$ are the Fourier transform and inverse Fourier transform, respectively. The fractional NLS also appears in the continuum limit of discrete models with long-range interactions (see [22]) and in the description of Bonson stars as well as in water wave dynamics (see [15]). Recently, an optical realization of the fractional Schrödinger equation was proposed by Longhi [28].

In this article, we consider the blow-up criteria and the strong instability of standing waves for the fractional Schrödinger-Poisson equation

$$
\begin{align*}
i \partial_{t} \psi-(-\Delta)^{s} \psi-\phi \psi+|\psi|^{p} \psi & =0, \quad(t, x) \in\left[0, T^{*}\right) \times \mathbb{R}^{3}, \\
(-\Delta)^{r} \phi & =|\psi|^{2}, \tag{1.2}
\end{align*}
$$

[^0]where $\psi:\left[0, T^{*}\right) \times \mathbb{R}^{3} \rightarrow \mathbb{C}$ is the complex valued function, $s, r \in(0,1)$, and $0<T^{*} \leq \infty, 0<p<\frac{4 s}{3-2 s}$. Under this assumption, $\phi$ can be expressed as
\[

$$
\begin{equation*}
\phi(x)=c_{r} \int_{\mathbb{R}^{3}} \frac{|\psi(y)|^{2}}{|x-y|^{3-2 r}} d y \tag{1.3}
\end{equation*}
$$

\]

which is called the $r$-Riesz potential, where

$$
c_{r}=\pi^{-3 / 2} 2^{-2 r} \frac{\Gamma\left(\frac{3}{2}-2 r\right)}{\Gamma(r)}
$$

In (1.3), and in the sequel, in we often omit the constant $c_{r}$ for convenience of notation. Substituting $\phi$ into 1.2 leads to the fractional Schrödinger equation

$$
\begin{gather*}
i \partial_{t} \psi-(-\Delta)^{s} \psi-\left(|x|^{-(3-2 r)} *|\psi|^{2}\right) \psi+|\psi|^{p} \psi=0, \quad(t, x) \in\left[0, T^{*}\right) \times \mathbb{R}^{3} \\
\psi(0, x)=\psi_{0}(x) \tag{1.4}
\end{gather*}
$$

where $\psi_{0} \in H^{s}$.
For the classical NLS, i.e., $s=1$, we have the Variance-Virial Law

$$
\begin{equation*}
\frac{1}{2} \frac{d}{d t} \int_{\mathbb{R}^{3}}|x|^{2}|\psi(t, x)|^{2} d x=2 \operatorname{Im} \int_{\mathbb{R}^{3}} \bar{\psi}(t, x) x \cdot \nabla \psi(t, x) d x \tag{1.5}
\end{equation*}
$$

provided that $\psi_{0} \in \Sigma:=\left\{v \in H^{1}: x v \in L^{2}\right\}$, where Im denotes the imaginary part. By using (1.5) and the virial identity, we can obtain the blow-up results for the classical NLS with negative energy $E\left(\psi_{0}\right)<0$ and finite variance [5]. However, this argument breaks down for $0<s<1$, since identity 1.5 fails in this case by the dimensional analysis. It turns out that the suitable generalization of the variance for the fractional NLS is

$$
\begin{equation*}
\mathcal{V}^{(s)}[\psi(t)]:=\int_{\mathbb{R}^{3}} \bar{\psi}(t, x) x \cdot(-\Delta)^{1-s} x \psi(t, x) d x=\left\|x(-\Delta)^{\frac{1-s}{2}} \psi(t)\right\|_{L^{2}}^{2} \tag{1.6}
\end{equation*}
$$

Given any sufficiently regular and spatially localized solution $\psi(t)$ of the free fractional Schrödinger equation $i \partial_{t} \psi=(-\Delta)^{s} \psi$, a calculation yields

$$
\begin{equation*}
\frac{1}{2} \frac{d}{d t} \mathcal{V}^{(s)}[\psi(t)]:=2 \operatorname{Im} \int_{\mathbb{R}^{3}} \bar{\psi}(t, x) x \cdot \nabla \psi(t, x) d x \tag{1.7}
\end{equation*}
$$

This idea has been successfully applied to prove the blow-up results for 1.1 with radial solutions and the Hartree-type nonlinearity $\left(|x|^{-\gamma} *|\psi|^{2}\right) \psi$ with $\gamma \geq 1$ in [6, 43]. But this method can not work due to the nontrivial error terms which seem very hard to control for the local nonlinearity $|\psi|^{p} \psi$. Boulenger et al. [3] applied the Balakrishman's formula

$$
\begin{equation*}
(-\Delta)^{s}=\frac{\sin \pi s}{\pi} \int_{0}^{\infty} m^{s-1} \frac{-\Delta}{-\Delta+m} d m \tag{1.8}
\end{equation*}
$$

and obtained the differential estimate

$$
\begin{aligned}
& \frac{d}{d t}\left(\operatorname{Im} \int_{\mathbb{R}^{3}} \bar{\psi}(t) \nabla \varphi_{R} \cdot \nabla \psi(t) d x\right) \\
& \leq 12 p E\left(\psi_{0}\right)-2 \delta\left\|(-\Delta)^{s / 2} \psi(t)\right\|_{L^{2}}^{2}+o_{R}(1)\left(1+\left\|(-\Delta)^{s / 2} \psi(t)\right\|_{L^{2}}^{p / s+}\right)
\end{aligned}
$$

where $\delta=3 p-2 s$. With the help of this key estimate, they proved the existence of radial blow-up $H^{s}$ solutions by applying the comparison theory.

However, to the best of our knowledge, there are no any blow-up results for (1.4) so far. In particular, equation (1.4) includes two classical nonlinearities, i.e., power-type $|\psi|^{p} \psi$ and Hartree-type $\left(|x|^{-(3-2 r)} *|\psi|^{2}\right) \psi$. The study of blow-up
solutions for $(1.4)$ is of particular challenge, because the methods for proving blowup results of (1.1) with power-type $|\psi|^{p} \psi$ or Hartree-type $\left(|x|^{-(3-2 r)} *|\psi|^{2}\right) \psi$ are usually different, so we should develop a new method when both nonlinearities appear simultaneously.

Inspired by the ideas in [9], we study the blow-up criteria for 1.4 . The difficulty is the presence of the fractional order Laplacian $(-\Delta)^{s}$. When $s=1$, we have

$$
\begin{equation*}
\frac{1}{2} \frac{d}{d t} \int_{\mathbb{R}^{3}} \varphi(x)|\psi(t, x)|^{2} d x=2 \operatorname{Im} \int_{\mathbb{R}^{3}} \bar{\psi}(t, x) \nabla \varphi(x) \cdot \nabla \psi(t, x) d x \tag{1.9}
\end{equation*}
$$

Using this identity, Du et al. 9 derived an $L^{2}$-estimate in the exterior ball. Thanks to this $L^{2}$-estimate and the virial estimates, they established the blow-up criteria for the classical NLS. In the case $s \in\left(\frac{1}{2}, 1\right)$, the identity 1.9 does not hold. However, by exploiting the idea in [3] and the use of the Balakrishman's formula (1.8), we can obtain the time derivative of the virial action. Thus, we can obtain the blow-up criteria for (1.4).

Theorem 1.1. Let $s \in(1 / 2,1)$ and $\psi_{0} \in H^{s}$ be the corresponding (not necessary radial) solution to 1.4 on the maximal time interval $\left[0, T^{*}\right)$. If there exists $\delta>0$ such that

$$
\begin{equation*}
\sup _{t \in\left[0, T^{*}\right)} Q(\psi(t)) \leq-\delta<0 \tag{1.10}
\end{equation*}
$$

where $Q(\psi(t))$ is defined by 1.14 . Then one of the following statements is true:

- $\psi(t)$ blows up in finite time, i.e. $T^{*}<+\infty$; or
- $\psi(t)$ blows up in infinite time and there exists a time sequence $\left(t_{n}\right)_{n \geq 1}$ such that $t_{n} \rightarrow+\infty$ and

$$
\begin{equation*}
\lim _{n \rightarrow \infty}\left\|(-\Delta)^{s / 2} \psi\left(t_{n}\right)\right\|_{L^{2}}=\infty \tag{1.11}
\end{equation*}
$$

Based on the blow-up criterion 1.10 , we will study the strong instability of standing waves of (1.4). The standing waves of $\sqrt{1.4}$ are solutions of the form $e^{i \omega t} u$, where $\omega \in \mathbb{R}$ is a frequency and $u \in H^{s} \backslash\{0\}$ is a nontrivial solution to the elliptic equation

$$
\begin{equation*}
(-\Delta)^{s} u+\omega u+\left(|x|^{-(3-2 r)} *|u|^{2}\right) u-|u|^{p} u=0 \tag{1.12}
\end{equation*}
$$

Note that 1.12 can be written as $S_{\omega}^{\prime}(u)=0$, where

$$
\begin{align*}
S_{\omega}(u): & =\frac{1}{2}\|u\|_{\dot{H}^{s}}^{2}+\frac{\omega}{2}\|u\|_{L^{2}}^{2}+\frac{1}{4} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|u|^{2}\right)|u|^{2} d x \\
& -\frac{1}{p+2}\|u\|_{L^{p+2}}^{p+2} \tag{1.13}
\end{align*}
$$

is the action functional. Then we define

$$
\begin{align*}
Q(u) & :=\left.\partial_{\lambda} S_{\omega}\left(u^{\lambda}\right)\right|_{\lambda=1} \\
& =s\|u\|_{\dot{H}^{s}}^{2}+\frac{3-2 r}{4} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|u|^{2}\right)|u|^{2} d x-\frac{3 p}{2(p+2)}\|u\|_{L^{p+2}}^{p+2} \tag{1.14}
\end{align*}
$$

with $u^{\lambda}(x):=\lambda^{3 / 2} u(\lambda x)$ and

$$
\begin{align*}
K_{\omega}(u): & (s+r)\left\langle S_{\omega}^{\prime}(u), u\right\rangle-I_{\omega}(u) \\
= & \frac{4 s+2 r-3}{2}\|u\|_{\dot{H}^{s}}^{2}+\frac{\omega(2 s+2 r-3)}{2}\|u\|_{L^{2}}^{2} \\
& +\frac{4 s+2 r-3}{4} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|u|^{2}\right)|u|^{2} d x  \tag{1.15}\\
& -\frac{(s+r)(p+2)-3}{p+2}\|u\|_{L^{p+2}}^{p+2},
\end{align*}
$$

where $I_{\omega}(u)$ denotes the Pohozaev identity related to 1.12 , see 2.3$)$.
The usual strategy to show the strong instability of standing waves of the classical NLS ( $\mathrm{s}=1$ ) is to establish the finite time blow-up by using the variational characterization of ground states as minimizers of the action functional and the virial identity. More specifically, the variational characterization of ground states by the manifold $\mathcal{N}:=\left\{v \in H^{1} \backslash\{0\}, Q(v)=0\right\}$ can imply the key estimate $Q(\psi(t)) \leq 2\left(S_{\omega}\left(\psi_{0}\right)-S_{\omega}(u)\right)$, where $u$ is the ground state solution. Then, it follows from the virial identity and the choice of initial data $\psi_{0}$ that

$$
\frac{d^{2}}{d t^{2}}\|x \psi(t)\|_{L^{2}}^{2}=8 Q(\psi(t)) \leq 16\left(S_{\omega}\left(\psi_{0}\right)-S_{\omega}(u)\right)<0
$$

for $t \in\left[0, T^{*}\right)$. This implies that the solution $\psi(t)$ blows up in a finite time. Thus, we can prove the strong instability of ground state standing waves [5, 26, 31, 32, 33, 37. However, in many cases, it is hard to obtain the variational characterization of ground states by the manifold $\mathcal{N}$. But we can obtain the variational characterization of ground states by the Nehari manifold and obtain the key estimate $Q(\psi(t)) \leq$ $2\left(S_{\omega}\left(\psi_{0}\right)-S_{\omega}(u)\right)$ [16, 17, 18, 27, 19, 29, 30, 34, 38, 41].

When $s=r=1$ and $p \in\{2 / 3\} \cup(1,4 / 3)$, Bellazzini and Siciliano [1] proved the existence of orbitally stable standing waves for (1.4). Kikuchi [25] showed the existence of standing waves for 1.4 with $s=r=1$ and $0<p<4$ and proved that the standing wave $e^{i \omega t} u$ is strongly unstable for all $\omega>0$ when $2 \leq p<4$. When $4 / 3<p<2$, it shows that there exists $\bar{\omega}>0$ such that the standing wave $e^{i \omega t} u$ is strongly unstable for all $\omega>\bar{\omega}$. In the $L^{2}$-supercritical case, i.e., $4 / 3<p<4$, Bellazzini et al. 2] improved the result of Kikuchi and proved that the standing wave $e^{i \omega t} u$ is strongly unstable for all $\omega>0$. When $4 / 3 \leq p<4$, Feng et al. [13] proved that the standing wave $e^{i \omega t} u$ is strongly unstable for all $\omega>0$.

Equation (1.4) with $p=4 s / 3$ is a class of nonlinear Schrödinger equations with combined $L^{2}$-critical and $L^{2}$-subcritical nonlinearities. When we try to study the variational characterization of ground states by the manifold $\mathcal{N}$, it is hard to obtain $S_{\omega}^{\prime}(u)=0$, see Lemma 5.4. Moreover, we find that the usual Nehari manifold is not a good choice in this case. Fortunately, we can obtain the variational characterization of ground states by the Nehari-Pohozaev manifold $\mathcal{N}_{1}:=\{u \in$ $\left.H^{s} \backslash\{0\}, K_{\omega}(u)=0\right\}$. Based on this variational characterization and a theoretical analysis, we can obtain the strong instability of standing waves for 1.4.

Theorem 1.2. Let $\omega>0, s \in(1 / 2,1), 2 s+2 r>3,4 s / 3 \leq p<4 s /(3-2 s)$ and $u$ be the ground state related to (1.12). Then the standing wave $\psi(t, x)=e^{i \omega t} u(x)$ is strongly unstable in the following sense: there exists $\left\{\psi_{0, n}\right\} \subset H^{s}$ such that $\psi_{0, n} \rightarrow u$ in $H^{s}$ as $n \rightarrow \infty$ and the corresponding solution $\psi_{n}$ of (1.4 with initial data $\psi_{0, n}$ blows up in finite time or infinite time for any $n \geq 1$.

This article is organized as follows. In Section 2, we present some useful lemmas such as the local well-posedness theory of (1.4), Brezis-Lieb's lemma, and the compactness lemma. In Section 3, we prove the localized virial estimates related to 1.4. In Sections 4 and 5, we prove Theorems 1.1 and 1.2 , respectively.

## 2. Preliminary lemmas

In this section, we recall some preliminary results that will be used later. Firstly, let us recall the local theory for the Cauchy problem (1.4). The local well-posedness for the fractional NLS in the energy space $H^{s}$ was studied by Hong and Sire in [20]. The proof is based on Strichartz's estimates and the contraction mapping argument. Note that for non-radial data, Strichartz's estimates have a loss of derivatives. Fortunately, this loss of derivatives can be compensated by using the Sobolev embedding. However, it leads to a weak local well-posedness in the energy space compared to the classical nonlinear Schrödinger equation. We refer the reader to [7. 20] for more details. We can remove the loss of derivatives in Strichartz's estimates by considering radially symmetric data. However, it needs a restriction on the validity of $s$, namely $\frac{3}{5} \leq s<1$.
Proposition 2.1. Let $3 / 5 \leq s<1,0<p<\frac{4 s}{3-2 s}$ and $\psi_{0} \in H^{s}$ be radial. Then there exists $T=T\left(\left\|\psi_{0}\right\|_{H^{s}}\right)$ such that (1.4) admits a unique solution $\psi \in$ $C\left([0, T], H^{s}\right)$. Let $\left[0, T^{*}\right)$ be the maximal time interval on which the solution $\psi$ is well-defined. If $T^{*}<\infty$, then $\|\psi(t)\|_{\dot{H}^{s}} \rightarrow \infty$ as $t \uparrow T^{*}$. Moreover, for all $0 \leq t<T^{*}$, the solution $\psi(t)$ satisfies the following conservation of mass and energy

$$
\begin{aligned}
\|\psi(t)\|_{L^{2}} & =\left\|\psi_{0}\right\|_{L^{2}} \\
E(\psi(t)) & =E\left(\psi_{0}\right)
\end{aligned}
$$

where

$$
\begin{align*}
E(\psi(t))= & \frac{1}{2}\|\psi(t)\|_{\dot{H}^{s}}^{2}+\frac{1}{4} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|\psi(t)|^{2}\right)(x)|\psi(t, x)|^{2} d x  \tag{2.1}\\
& -\frac{1}{p+2}\|\psi(t)\|_{L^{p+2}}^{p+2}
\end{align*}
$$

In this article, we use the so called Brezis-Lieb's lemma [4].
Lemma 2.2. Let $0<p<\infty$. Suppose that $u_{n} \rightarrow u$ almost everywhere and $\left\{u_{n}\right\}$ is a bounded sequence in $L^{p}$. Then

$$
\lim _{n \rightarrow \infty}\left(\left\|u_{n}\right\|_{L^{p}}^{p}-\left\|u_{n}-u\right\|_{L^{p}}^{p}-\|u\|_{L^{p}}^{p}\right)=0
$$

Lemma 2.3 (40). Let $u \in H^{s}$ and $2 s+2 r>3$. Suppose that $u_{n} \rightharpoonup u$ in $H^{s}$ and $u_{n} \rightarrow u$ a.e. in $\mathbb{R}^{3}$. Then

$$
\begin{aligned}
\int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *\left|u_{n}\right|^{2}\right)\left|u_{n}\right|^{2} d x= & \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *\left|u_{n}-u\right|^{2}\right)\left|u_{n}-u\right|^{2} d x \\
& +\int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|u|^{2}\right)|u|^{2} d x+\circ(1)
\end{aligned}
$$

The following compactness lemma is vital in our discussions [8, 10].

Lemma 2.4. Let $0<p<\frac{4 s}{3-2 s}$ and $\left\{u_{n}\right\}$ be a bounded sequence in $H^{s}$ such that

$$
\limsup _{n \rightarrow \infty}\left\|u_{n}\right\|_{\dot{H}^{s}} \leq M, \quad \liminf _{n \rightarrow \infty}\left\|u_{n}\right\|_{L^{p+2}} \geq m
$$

Then there exist a sequence $\left(x_{n}\right)_{n \geq 1}$ in $\mathbb{R}^{3}$ and $u \in H^{s} \backslash\{0\}$ such that up to $a$ subsequence,

$$
u_{n}\left(\cdot+x_{n}\right) \rightharpoonup u \quad \text { weakly in } H^{s} .
$$

Finally, we recall the Pohozaev identity related to (1.12) 40.
Lemma 2.5. If $u \in H^{s}$ satisfies equation (1.12), then it holds

$$
\begin{equation*}
\|u\|_{\dot{H}^{s}}^{2}+\omega\|u\|_{L^{2}}^{2}+\int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|u|^{2}\right)|u|^{2} d x-\|u\|_{L^{p+2}}^{p+2}=0 \tag{2.2}
\end{equation*}
$$

and

$$
\begin{align*}
I_{\omega}(u):= & \frac{3-2 s}{2}\|u\|_{\dot{H}^{s}}^{2}+\frac{3 \omega}{2}\|u\|_{L^{2}}^{2}+\frac{2 r+3}{4} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|u|^{2}\right)|u|^{2} d x  \tag{2.3}\\
& -\frac{3}{p+2}\|u\|_{L^{p+2}}^{p+2}=0
\end{align*}
$$

## 3. Localized virial estimates

In this section, we prove some localized virial estimates related to 1.4 . Let us recall some useful results in [3].
Lemma 3.1 ([3]). Suppose $\varphi: \mathbb{R}^{3} \rightarrow \mathbb{R}$ is such that $\nabla \varphi \in W^{1, \infty}$. Then for all $u \in H^{1 / 2}$, it holds

$$
\left|\int_{\mathbb{R}^{3}} \bar{u}(x) \nabla \varphi(x) \cdot \nabla u(x) d x\right| \leq C\|\nabla \varphi\|_{W^{1, \infty}}\left(\left\||\nabla|^{1 / 2} u\right\|_{L^{2}}^{2}+\|u\|_{L^{2}}\left\||\nabla|^{1 / 2} u\right\|_{L^{2}}\right)
$$

for some constant $C>0$.
To study the localized virial estimates for (1.4), we introduce the auxiliary function

$$
\begin{equation*}
u_{m}(x):=c_{s} \frac{1}{-\Delta+m} u(x)=c_{s} \mathcal{F}^{-1}\left(\frac{\hat{u}(\xi)}{|\xi|^{2}+m}\right), \quad m>0 \tag{3.1}
\end{equation*}
$$

where

$$
c_{s}:=\sqrt{\frac{\sin \pi s}{\pi}}
$$

Lemma $3.2([3])$. Let $s \in(0,1)$ and $\varphi: \mathbb{R}^{3} \rightarrow \mathbb{R}$ with $\Delta \varphi \in W^{2, \infty}$. Then for all $u \in L^{2}$ it holds

$$
\left.\left|\int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}}\left(\Delta^{2} \varphi\right)\right| u_{m}\right|^{2} d x d m \mid \leq C\left\|\Delta^{2} \varphi\right\|_{L^{\infty}}^{s}\|\Delta \varphi\|_{L^{\infty}}^{1-s}\|u\|_{L^{2}}^{2}
$$

for some constant $C>0$ dependent only on $s$.
We refer the reader to [3, Appendix A] for the proof of Lemmas 3.1 and 3.2. Given that

$$
\frac{\sin \pi s}{\pi} \int_{0}^{\infty} \frac{m^{s}}{\left(|\xi|^{2}+m\right)^{2}} d m=s|\xi|^{2 s-2}
$$

Plancherel's and Fubini's theorems imply that

$$
\begin{align*}
\int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}}\left|\nabla u_{m}\right|^{2} d x d m & =\int_{\mathbb{R}^{3}}\left(\frac{\sin \pi s}{\pi} \int_{0}^{\infty} \frac{m^{s} d m}{\left(|\xi|^{2}+m\right)^{2}}\right)|\xi|^{2}|\hat{u}(\xi)|^{2} d \xi  \tag{3.2}\\
& =\int_{\mathbb{R}^{3}}\left(s|\xi|^{2 s-2}\right)|\xi|^{2}|\hat{u}(\xi)|^{2} d \xi=s\left\|(-\Delta)^{s / 2} u\right\|_{L^{2}}^{2}
\end{align*}
$$

for any $u \in \dot{H}^{s}$.
Lemma 3.3. Let $s \in(1 / 2,1)$ and $\varphi: \mathbb{R}^{3} \rightarrow \mathbb{R}$ be such that $\nabla \varphi \in W^{1, \infty}$. Then for any $u \in L^{2}$ it holds

$$
\left.\left|\int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}}(\Delta \varphi)\right| u_{m}\right|^{2} d x d m \mid \leq C\|\Delta \varphi\|_{L^{\infty}}^{2 s-1}\|\nabla \varphi\|_{L^{\infty}}^{2-2 s}\|u\|_{L^{2}}^{2}
$$

for some constant $C>0$ dependent only on $s$.
Proof. The idea is essentially similar to [3, Lemma A.2]. For the reader's convenience, we just present the outline of our proof. Splitting $m$-integral into $\int_{0}^{\rho} \ldots$ and $\int_{\rho}^{\infty} \ldots$ with $\rho>0$ to be chosen later. For the first term, we use integration by parts and Hölder's inequality to have

$$
\begin{aligned}
\left.\left|\int_{0}^{\rho} m^{s} \int_{\mathbb{R}^{3}}(\Delta \varphi)\right| u_{m}\right|^{2} d x d m \mid & =\left|\int_{0}^{\rho} m^{s} \int_{\mathbb{R}^{3}} \nabla \varphi \cdot\left(\nabla u_{m} \bar{u}_{m}+u_{m} \nabla \bar{u}_{m}\right) d x d m\right| \\
& =\|\nabla \varphi\|_{L^{\infty}} \int_{0}^{\rho} m^{s}\left\|\nabla u_{m}\right\|_{L^{2}}\left\|u_{m}\right\|_{L^{2}} d m \\
& =\|\nabla \varphi\|_{L^{\infty}}\|u\|_{L^{2}}^{2}\left(\int_{0}^{\rho} m^{s-3 / 2} d m\right) \\
& \leq C \rho^{s-1 / 2}\|\nabla \varphi\|_{L^{\infty}}\|u\|_{L^{2}}^{2} .
\end{aligned}
$$

Here we use the fact $\left\|\nabla u_{m}\right\|_{L^{2}} \leq C m^{-1 / 2}\|u\|_{L^{2}}$ and $\left\|u_{m}\right\|_{L^{2}} \leq C m^{-1}\|u\|_{L^{2}}$ which follows from the definition of $u_{m}$. For the second term, we have

$$
\begin{aligned}
\left.\left|\int_{\rho}^{\infty} m^{s} \int_{\mathbb{R}^{3}}(\Delta \varphi)\right| u_{m}\right|^{2} d x d m \mid & \leq C\|\Delta \varphi\|_{L^{\infty}}\left(\int_{\rho}^{\infty} m^{s}\left\|u_{m}\right\|_{L^{2}}^{2} d m\right) \\
& \leq C\|\Delta \varphi\|_{L^{\infty}}\|u\|_{L^{2}}^{2}\left(\int_{\rho}^{\infty} m^{s-2} d m\right) \\
& \leq C \rho^{s-1}\|\Delta \varphi\|_{L^{\infty}}\|u\|_{L^{2}}^{2} .
\end{aligned}
$$

Combining two terms yields

$$
\left.\left|\int_{0}^{\infty} \int_{\mathbb{R}^{3}}(\Delta \varphi)\right| u_{m}\right|^{2} d x d m \mid \leq C\left(\rho^{s-1 / 2}\|\nabla \varphi\|_{L^{\infty}}+\rho^{s-1}\|\Delta \varphi\|_{L^{\infty}}\right)\|u\|_{L^{2}}^{2}
$$

for arbitrary $\rho>0$. Minimizing the right hand side with respect to $\rho$, i.e. choosing $\rho=\left(\frac{(1-s)\|\Delta \varphi\|_{L^{\infty}}}{(s-1 / 2)\|\nabla \varphi\|_{L^{\infty}}}\right)^{2}$, we obtain the desired result.

By the same argument as in Lemma 3.3 and Lemma 3.1. we obtain the following result.
Lemma 3.4. Let $s \in(1 / 2,1)$ and $\varphi: \mathbb{R}^{3} \rightarrow \mathbb{R}$ be such that $\nabla \varphi \in W^{1, \infty}$. Then for any $u \in H^{1 / 2}$ we have

$$
\left|\int_{0}^{\infty} \int_{\mathbb{R}^{3}} \bar{u}_{m} \nabla \varphi \cdot \nabla u_{m} d x d m\right| \leq C\|\nabla \varphi\|_{W^{1, \infty}}\|u\|_{H^{1 / 2}}^{2}
$$

for some constant $C>0$.

Let $1 / 2<s<1$ and $\varphi: \mathbb{R}^{3} \rightarrow \mathbb{R}$ be such that $\varphi \in W^{2, \infty}$. Assume that $\psi \in C\left(\left[0, T^{*}\right), H^{s}\right)$ is a solution to 1.4$]$. We define the localized virial action of $\psi$ associated to $\varphi$ by

$$
\mathcal{V}_{\varphi}[\psi(t)]:=\int_{\mathbb{R}^{3}} \varphi(x)|\psi(t, x)|^{2} d x
$$

Lemma 3.5 (Virial identity). Let $s \in(1 / 2,1)$ and $\varphi: \mathbb{R}^{3} \rightarrow \mathbb{R}$ be such that $\varphi \in W^{2, \infty}$. Assume that $\psi \in C\left(\left[0, T^{*}\right), H^{s}\right)$ is a solution to (1.4). Then for any $t \in\left[0, T^{*}\right)$ we have

$$
\begin{aligned}
& \frac{d}{d t} \mathcal{V}_{\varphi}[\psi(t)] \\
& =-i \int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}}(\Delta \varphi)\left|\psi_{m}(t)\right|^{2} d x d m-2 i \int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}} \bar{\psi}_{m}(t) \nabla \varphi \cdot \nabla \psi_{m}(t) d x d m
\end{aligned}
$$

where $\psi_{m}(t)=c_{s}(-\Delta+m)^{-1} \psi(t)$.
Proof. It suffices to prove Lemma 3.5 for $\psi(t) \in C_{0}^{\infty}\left(\mathbb{R}^{3}\right)$. The general case follows by an approximation argument. We write

$$
\mathcal{V}_{\varphi}[\psi(t)]=\langle\psi(t), \varphi \psi(t)\rangle
$$

where $\langle\cdot, \cdot\rangle$ is the scalar product in $L^{2}$. Since $\psi(t)$ satisfies 1.4 , it is easy to see that

$$
\frac{d}{d t} \mathcal{V}_{\varphi}[\psi(t)]=i\left\langle\psi(t),\left[(-\Delta)^{s}, \varphi\right] \psi(t)\right\rangle
$$

where $[X, Y]=X Y-Y X$ is the commutator of $X$ and $Y$. To study $\left[(-\Delta)^{s}, \varphi\right]$, we recall the Balakrishman's formula

$$
(-\Delta)^{s}=\frac{\sin \pi s}{\pi} \int_{0}^{\infty} m^{s-1} \frac{-\Delta}{-\Delta+m} d m
$$

Using the fact that for operators $A \geq 0, B$ with $m>0$ being any positive real number

$$
\left[\frac{A}{A+m}, B\right]=\left[1-\frac{m}{A+m}, B\right]=-m\left[\frac{1}{A+m}, B\right]=m \frac{1}{A+m}[A, B] \frac{1}{A+m}
$$

and letting $A=(-\Delta)^{s}, B=\varphi$ and using the Balakrishman's formula, we have

$$
\begin{aligned}
{\left[(-\Delta)^{s}, \varphi\right] } & =\frac{\sin \pi s}{\pi} \int_{0}^{\infty} m^{s}\left[\frac{-\Delta}{-\Delta+m}, \varphi\right] d m \\
& =\frac{\sin \pi s}{\pi} \int_{0}^{\infty} m^{s} \frac{1}{-\Delta+m}[-\Delta, \varphi] \frac{1}{-\Delta+m} d m
\end{aligned}
$$

Thus we obtain

$$
\begin{aligned}
& \left\langle\psi(t),\left[(-\Delta)^{s}, \varphi\right] \psi(t)\right\rangle \\
& =\left\langle\psi(t),\left(\frac{\sin \pi s}{\pi} \int_{0}^{\infty} m^{s} \frac{1}{-\Delta+m}[-\Delta, \varphi] \frac{1}{-\Delta+m} d m\right) \psi(t)\right\rangle \\
& =c_{s}^{2} \int_{0}^{\infty} m^{s}\left\langle\psi(t), \frac{1}{-\Delta+m}[-\Delta, \varphi] \frac{1}{-\Delta+m} \psi(t)\right\rangle d m \\
& =\int_{0}^{\infty} m^{s}\left\langle c_{s}(-\Delta+m)^{-1} \psi(t),[-\Delta, \varphi] c_{s}(-\Delta+m)^{-1} \psi(t)\right\rangle d m \\
& =\int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}} \bar{\psi}_{m}(t)\left(-\Delta \varphi \psi_{m}(t)-2 \nabla \varphi \cdot \nabla \psi_{m}(t)\right) d x d m
\end{aligned}
$$

$$
=\int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}}\left((-\Delta \varphi)\left|\psi_{m}(t)\right|^{2}-2 \bar{\psi}_{m}(t) \nabla \varphi \cdot \nabla u_{m}(t)\right) d x d m
$$

A direct consequence of Lemmas 3.3 and 3.4 is the following estimate.
Lemma 3.6. Let $s \in(1 / 2,1)$ and $\varphi: \mathbb{R}^{3} \rightarrow \mathbb{R}$ be such that $\varphi \in W^{2, \infty}$. Assume that $\psi \in C\left(\left[0, T^{*}\right), H^{s}\right)$ is a solution to 1.4$)$. Then for any $t \in\left[0, T^{*}\right)$ we have

$$
\left|\frac{d}{d t} \mathcal{V}_{\varphi}[\psi(t)]\right| \leq C\|\nabla \varphi\|_{W^{1, \infty}}\|\psi(t)\|_{H^{s}}^{2}
$$

for some constant $C>0$ dependent only on $s$.
We next define the localized Morawetz action of $\psi$ associated to $\varphi$ by

$$
\begin{equation*}
\mathcal{M}_{\varphi}[\psi(t)]:=2 \operatorname{Im} \int_{\mathbb{R}^{3}} \bar{\psi}(t, x) \nabla \varphi(x) \cdot \nabla \psi(t, x) d x \tag{3.3}
\end{equation*}
$$

By Lemma 3.1, we obtain the bound

$$
\left|\mathcal{M}_{\varphi}[\psi(t)]\right| \leq C\left(\|\nabla \varphi\|_{L^{\infty}},\|\Delta \varphi\|_{L^{\infty}}\right)\|\psi(t)\|_{H^{1 / 2}}^{2}
$$

Hence the quantity $\mathcal{M}_{\varphi}[\psi(t)]$ is well-defined, given $\psi(t) \in H^{s}\left(\mathbb{R}^{3}\right)$ with $s>1 / 2$.
Lemma 3.7 (Morawetz identity). Let $s \in(1 / 2,1)$ and $\varphi: \mathbb{R}^{3} \rightarrow \mathbb{R}$ be such that $\nabla \varphi \in W^{3, \infty}$. Assume that $\psi \in C\left(\left[0, T^{*}\right), H^{s}\right)$ is a solution to 1.4). Then for each $t \in\left[0, T^{*}\right)$ we have

$$
\begin{align*}
& \frac{d}{d t} \mathcal{M}_{\varphi}[\psi(t)] \\
& =\int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}}\left\{4 \overline{\partial_{k} \psi_{m}(t)}\left(\partial_{k l}^{2} \varphi\right) \partial_{l} \psi_{m}(t)-\left(\Delta^{2} \varphi\right)\left|\psi_{m}(t)\right|^{2}\right\} d x d m \\
& \quad+(3-2 r) \int_{\mathbb{R}^{3}} \int_{\mathbb{R}^{3}} \frac{|u(t, x)|^{2}|u(t, y)|^{2}(x-y) \cdot(\nabla \varphi(x)-\nabla \varphi(y))}{|x-y|^{5-2 r}} d x d y  \tag{3.4}\\
& \quad-\frac{2 p}{p+2} \int_{\mathbb{R}^{3}} \Delta \varphi|\psi(t)|^{p+2} d x
\end{align*}
$$

where $\psi_{m}(t)=\psi_{m}(t, x)$ is defined in 3.1.
Proof. Integration by parts yields

$$
\begin{aligned}
& \left\langle u(t),\left[-\left(|x|^{-(3-2 r)} *|u(t)|^{2}\right), i \Gamma_{\varphi}\right] u(t)\right\rangle \\
& =-\left\langle u(t),\left[\left(|x|^{-(3-2 r)} *|u(t)|^{2}\right), \nabla \varphi \cdot \nabla+\nabla \cdot \nabla \varphi\right] u(t)\right\rangle \\
& =2 \int_{\mathbb{R}^{3}} \nabla \varphi \cdot \nabla\left(|x|^{-(3-2 r)} *|u(t)|^{2}\right)|u(t)|^{2} d x \\
& =-(3-2 r) \int_{\mathbb{R}^{3}} \int_{\mathbb{R}^{3}} \frac{|u(t, x)|^{2}|u(t, y)|^{2}(x-y) \cdot(\nabla \varphi(x)-\nabla \varphi(y))}{|x-y|^{5-2 r}} d x d y .
\end{aligned}
$$

The rest proof is similar to [3, Lemma 2.1], so we omit the details.
4. BLOW-UP CRITERIA FOR 1.4

Lemma 4.1. Let $\eta>0, R>1$ and the solution $\psi(t)$ of 1.4 satisfy

$$
\begin{equation*}
C_{1}:=\sup _{t \in[0,+\infty)}\|\psi(t)\|_{H^{s}}<\infty \tag{4.1}
\end{equation*}
$$

Then there exists a constant $C>0$ independent of $R$ and $C_{1}$ such that

$$
\int_{|x| \geq R}|\psi(t, x)|^{2} d x \leq \eta+o_{R}(1)
$$

for all $t \in\left[0, T_{0}\right]$ with $T_{0}:=\frac{\eta R}{C C_{1}^{2}}$.
Proof. Let us now introduce $\theta:[0, \infty) \rightarrow[0,1]$ a smooth function satisfying

$$
\theta(r)= \begin{cases}0 & \text { if } 0 \leq r \leq 1 / 2 \\ 1 & \text { if } r \geq 1\end{cases}
$$

For $R>1$, we denote the radial function

$$
\phi_{R}(x):=\theta(r / R), \quad r=|x|
$$

We have

$$
\nabla \phi_{R}(x)=\frac{x}{r R} \theta^{\prime}(r / R), \quad \Delta \phi_{R}(x)=\frac{1}{R^{2}} \theta^{\prime \prime}(r / R)+\frac{2}{r R} \theta^{\prime}(r / R)
$$

In particular, we have

$$
\begin{equation*}
\left\|\nabla \phi_{R}\right\|_{W^{1, \infty}} \sim\left\|\nabla \phi_{R}\right\|_{L^{\infty}}+\left\|\Delta \phi_{R}\right\|_{L^{\infty}} \leq C R^{-1} \tag{4.2}
\end{equation*}
$$

We define the localized virial potential as

$$
\mathcal{V}_{\phi_{R}}[\psi(t)]:=\int_{\mathbb{R}^{3}} \phi_{R}(x)|\psi(t, x)|^{2} d x
$$

We have

$$
\begin{aligned}
\mathcal{V}_{\phi_{R}}[\psi(t)] & =\mathcal{V}_{\phi_{R}}\left[\psi_{0}\right]+\int_{0}^{t} \frac{d}{d \tau} \mathcal{V}_{\phi_{R}}[\psi(\tau)] d \tau \\
& \leq \mathcal{V}_{\phi_{R}}\left[\psi_{0}\right]+\left(\sup _{\tau \in[0, t]}\left|\frac{d}{d \tau} \mathcal{V}_{\phi_{R}}[\psi(\tau)]\right|\right) t
\end{aligned}
$$

By Lemma 3.6, 4.1) and 4.2, we obtain

$$
\sup _{\tau \in[0, t]}\left|\frac{d}{d \tau} \mathcal{V}_{\phi_{R}}[\psi(\tau)]\right| \leq C\left\|\nabla \phi_{R}\right\|_{W^{1, \infty}} \sup _{\tau \in[0, t]}\|\psi(\tau)\|_{H^{s}}^{2} \leq C C_{1}^{2} R^{-1}
$$

for some constant $C>0$ independent of $R$ and $C_{1}$. Therefore,

$$
\mathcal{V}_{\phi_{R}}[\psi(t)] \leq \mathcal{V}_{\phi_{R}}\left[\psi_{0}\right]+C C_{1}^{2} R^{-1} t
$$

for all $t \geq 0$. By the choice of $\theta$ and the conservation of mass, we have

$$
\mathcal{V}_{\phi_{R}}\left[\psi_{0}\right]=\int_{\mathbb{R}^{3}} \phi_{R}(x)\left|\psi_{0}(x)\right|^{2} d x \leq \int_{|x|>R / 2}\left|\psi_{0}(x)\right|^{2} d x \rightarrow 0
$$

as $R \rightarrow \infty$ or $\mathcal{V}_{\phi_{R}}\left[\psi_{0}\right]=o_{R}(1)$. On the other hand,

$$
\int_{|x| \geq R}|\psi(t, x)|^{2} d x \leq \mathcal{V}_{\phi_{R}}[\psi(t)]
$$

Combing the above estimates, we arrive at the desired result.

Proof of Theorem 1.1. If $T^{*}<+\infty$, then the proof is done. If $T^{*}=+\infty$, then we need to show 1.11 . Let $\varphi: \mathbb{R}^{3} \rightarrow \mathbb{R}$ be such that $\nabla \varphi \in W^{3, \infty}$. In addition, we assume that $\varphi=\varphi(r)$ is radial and satisfies

$$
\varphi(r)= \begin{cases}\frac{r^{2}}{2} & \text { for } r \leq 1 \\ \text { const. } & \text { for } r \geq 10\end{cases}
$$

and $\varphi^{\prime \prime}(r) \leq 1$ for $r \geq 0$. Given $R>0$, we define the rescaled function $\varphi_{R}: \mathbb{R}^{3} \rightarrow \mathbb{R}$ by

$$
\begin{equation*}
\varphi_{R}(r):=R^{2} \varphi\left(\frac{x}{R}\right) \tag{4.3}
\end{equation*}
$$

We readily verify the inequalities

$$
1-\varphi_{R}^{\prime \prime}(r) \geq 0, \quad 1-\frac{\varphi_{R}^{\prime}(r)}{r} \geq 0, \quad 3-\Delta \varphi_{R}(x) \geq 0
$$

for all $r \geq 0$ and all $x \in \mathbb{R}^{3}$. It is easy to see that

$$
\left\|\nabla^{k} \varphi_{R}\right\|_{L^{\infty}} \leq R^{2-k}, \quad k=0, \ldots, 4
$$

and

$$
\operatorname{supt}\left(\nabla^{k} \varphi_{R}\right) \subset \begin{cases}\{|x| \leq 10 R\} & \text { for } k=1,2 \\ \{R \leq|x| \leq 10 R\} & \text { for } k=3,4\end{cases}
$$

Applying Lemma 3.7. we have

$$
\begin{align*}
& \frac{d}{d t} \mathcal{M}_{\varphi_{R}}[\psi(t)] \\
& =\int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}}\left\{4 \overline{\partial_{k} \psi_{m}(t)}\left(\partial_{k l}^{2} \varphi_{R}\right) \partial_{l} \psi_{m}(t)-\left(\Delta^{2} \varphi_{R}\right)\left|\psi_{m}(t)\right|^{2}\right\} d x d m \\
& \quad-\frac{2 p}{p+2} \int_{\mathbb{R}^{3}} \Delta \varphi_{R}|\psi(t)|^{p+2} d x  \tag{4.4}\\
& \quad+(3-2 r) \int_{\mathbb{R}^{3}} \int_{\mathbb{R}^{3}} \frac{|\psi(t, x)|^{2}|\psi(t, y)|^{2}(x-y) \cdot\left(\nabla \varphi_{R}(x)-\nabla \varphi_{R}(y)\right)}{|x-y|^{5-2 r}} d x d y
\end{align*}
$$

where $\psi_{m}(t)=\psi_{m}(t, x)$ is defined in 3.1). Since $\operatorname{supt}\left(\Delta^{2} \varphi_{R}\right) \subset\{|x| \geq R\}$, by Lemma 3.2, we have

$$
\begin{align*}
\left.\left|\int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}}\left(\Delta^{2} \varphi_{R}\right)\right| \psi_{m}(t)\right|^{2} d x d m \mid & \leq C\left\|\Delta^{2} \varphi_{R}\right\|_{L^{\infty}}^{s}\left\|\Delta \varphi_{R}\right\|_{L^{\infty}}^{1-s}\|\psi(t)\|_{L^{2}(|x| \geq R)}^{2} \\
& \leq C R^{-2 s}\|\psi(t)\|_{L^{2}(|x| \geq R)}^{2} \tag{4.5}
\end{align*}
$$

Since $\varphi_{R}$ is radial, we use

$$
\partial_{j k}^{2}=\left(\frac{\delta_{j k}}{r}-\frac{x_{j} x_{k}}{r^{3}}\right) \partial_{r}+\frac{x_{j} x_{k}}{r^{2}} \partial_{r}^{2}
$$

to deduce

$$
\begin{aligned}
& \int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}} \overline{\partial_{k} \psi_{m}(t)}\left(\partial_{j k}^{2} \varphi_{R}\right) \partial_{l} \psi_{m}(t) d x d m \\
& =\int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}} \frac{\varphi_{R}^{\prime}}{r}\left|\nabla \psi_{m}(t)\right|^{2} d x d m \\
& \quad+\int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}}\left(\frac{\varphi_{R}^{\prime \prime}}{r^{2}}-\frac{\varphi_{R}^{\prime}}{r^{3}}\right)\left|x \cdot \nabla \psi_{m}(t)\right|^{2} d x d m
\end{aligned}
$$

Using (3.2) leads to

$$
\begin{aligned}
& \int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}} \frac{\varphi_{R}^{\prime}}{r}\left|\nabla \psi_{m}(t)\right|^{2} d x d m \\
& =s\left\|(-\Delta)^{s / 2} \psi(t)\right\|_{L^{2}}^{2}+\int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}}\left(\frac{\varphi_{R}^{\prime}}{r}-1\right)\left|\nabla \psi_{m}(t)\right|^{2} d x d m
\end{aligned}
$$

Since $\varphi_{R}^{\prime \prime} \leq 1$, the Cauchy-Schwarz inequality implies

$$
\begin{aligned}
& \int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}}\left(\frac{\varphi_{R}^{\prime}}{r}-1\right)\left|\nabla \psi_{m}(t)\right|^{2} d x d m \\
& +\int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}}\left(\varphi_{R}^{\prime \prime}-\frac{\varphi_{R}^{\prime}}{r}\right) \frac{\left|x \cdot \nabla \psi_{m}(t)\right|^{2}}{r^{2}} d x d m \leq 0 .
\end{aligned}
$$

Thus, we have

$$
\begin{equation*}
4 \int_{0}^{\infty} m^{s} \int_{\mathbb{R}^{3}} \overline{\partial_{k} \psi_{m}(t)}\left(\partial_{j k}^{2} \varphi_{R}\right) \partial_{l} \psi_{m}(t) d x d m \leq 4 s\left\|(-\Delta)^{s / 2} \psi(t)\right\|_{L^{2}}^{2} \tag{4.6}
\end{equation*}
$$

Note that
$-\frac{2 p}{p+2} \int_{\mathbb{R}^{3}} \Delta \varphi_{R}|\psi(t)|^{p+2} d x=-\frac{6 p}{p+2}\|\psi(t)\|_{L^{p+2}}^{p+2}+\frac{2 p}{p+2} \int_{\mathbb{R}^{3}}\left(3-\Delta \varphi_{R}\right)|\psi(t)|^{p+2} d x$.
Since $\operatorname{supt}\left(3-\Delta \varphi_{R}\right) \subset\{|x| \geq R\}$ and $\left\|3-\Delta \varphi_{R}\right\|_{L^{\infty}} \leq C$, we have

$$
\begin{aligned}
\int_{\mathbb{R}^{3}}\left(3-\Delta \varphi_{R}\right)|\psi(t)|^{p+2} d x & \leq C \int_{|x| \geq R}|\psi(t)|^{p+2} d x \\
& \leq C\|\psi(t)\|_{L^{\frac{3 p}{2 s}}} \quad \frac{6}{\frac{3}{2}^{3-2 s}}(|x| \geq R)
\end{aligned}\|\psi\|_{L^{2}(|x| \geq R)}^{\frac{4 s-(3-2 s) p}{2 s}}
$$

Thus we obtain

$$
\begin{equation*}
-\frac{2 p}{p+2} \int_{\mathbb{R}^{3}} \Delta \varphi_{R}|\psi(t)|^{p+2} d x \leq-\frac{6 p}{p+2}\|\psi(t)\|_{L^{p+2}}^{p+2}+C C_{1}^{\frac{3 p}{2 s}}\|\psi(t)\|_{L^{2}(|x| \geq R)}^{\frac{4 s-(3-2 s) p}{2 s}} \tag{4.7}
\end{equation*}
$$

We denote the last term in 4.4 by $\mathcal{T}$. We have

$$
\mathcal{T}=(3-2 r) \int_{\mathbb{R}^{3}} \int_{\mathbb{R}^{3}}(x-y) \cdot\left(\nabla \varphi_{R}(x)-\nabla \varphi_{R}(y)\right) \frac{|\psi(t, x)|^{2}|\psi(t, y)|^{2}}{|x-y|^{5-2 r}} d x d y
$$

By using

$$
\operatorname{supt}\left(|x-y|^{2}-(x-y) \cdot\left(\nabla \varphi_{R}(x)-\nabla \varphi_{R}(y)\right)\right) \subset\{|x| \geq R\} \cup\{|y| \geq R\}
$$

in the region $\{|x| \geq R\}$ we obtain

$$
\left||x-y|^{2}-(x-y) \cdot\left(\nabla \varphi_{R}(x)-\nabla \varphi_{R}(y)\right)\right| \leq C|x-y|^{2} .
$$

Thus, we obtain

$$
\begin{aligned}
& \left|\int_{|x| \geq R} \int_{\mathbb{R}^{3}}\left[|x-y|^{2}-(x-y) \cdot\left(\nabla \varphi_{R}(x)-\nabla \varphi_{R}(y)\right)\right] \frac{|u(t, x)|^{2}|\psi(t, y)|^{2 p_{2}}}{|x-y|^{5-2 r}} d x d y\right| \\
& \leq C \int_{|x| \geq R}\left(|x|^{-(3-2 r)} *|\psi(t)|^{2}\right)|\psi(t)|^{2} d x .
\end{aligned}
$$

To estimate this term, we deduce from the Sobolev embedding that

$$
\begin{equation*}
\|\psi(t)\|_{L^{\frac{12}{3+2 r}}}^{2} \leq C\|\psi(t)\|_{L^{2}}^{\frac{4 s+2 r-3}{2 s}}\|\psi(t)\|_{L^{\frac{3-2 r}{3-2 s}}}^{\frac{3-2 r}{2 s}} \leq C\left\|(-\Delta)^{s / 2} \psi(t)\right\|_{L^{\frac{3-2 r}{2 s}}}^{\frac{\frac{3-}{2 s}}{3-2}} \tag{4.8}
\end{equation*}
$$

Thus, it follows from the Hardy-Littlewood-Sobolev inequality and the conservation of mass that

$$
\begin{aligned}
& \int_{|x| \geq R}\left(|x|^{-(3-2 r)} *|\psi(t)|^{2}\right)|\psi(t)|^{2} d x \\
& \leq C\left\||x|^{-(3-2 r)} *|\psi(t)|^{2}\right\|_{L^{\frac{6}{3-2 r}}(|x| \geq R)}\left\||\psi(t)|^{2}\right\|_{L^{\frac{6}{3+2 r}}(|x| \geq R)} \\
& \leq C\|\psi(t)\|_{L^{\frac{12}{332 r}}}^{2}\|\psi(t)\|_{L^{\frac{12}{3+2 r}}}^{2}(|x| \geq R) \\
& \leq C\|\psi(t)\|_{H^{\frac{3-2 r}{s}}}^{\frac{3}{s}}\|\psi(t)\|_{L^{2}(|x| \geq R)}^{2 s-3} \\
& \leq C C_{1}^{\frac{4-2 r}{s}}\|\psi(t)\|_{L^{2}(|x| \geq R)}^{\frac{4 s+2 r-3}{2 s}} .
\end{aligned}
$$

We can derive an estimate in the region $\{|y| \geq R\}$ too. Similarly, we can obtain

$$
\begin{equation*}
\mathcal{T} \leq(3-2 r) \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|\psi(t)|^{2}\right)|\psi(t)|^{2} d x+C C_{1}^{\frac{3-2 r}{s}}\|\psi(t)\|_{L^{2}(|x| \geq R)}^{\frac{4 s+2 r-3}{2 s}} \tag{4.9}
\end{equation*}
$$

By using (4.5)-(4.9), we obtain

$$
\begin{align*}
& \frac{d}{d t} \mathcal{M}_{\varphi_{R}}[\psi(t)] \\
& \leq 4 s\left\|(-\Delta)^{s / 2} \psi(t)\right\|_{L^{2}}^{2}+C R^{-2 s}\|\psi(t)\|_{L^{2}(|x| \geq R)}^{2} \\
& \quad+(3-2 r) \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|\psi(t)|^{2}\right)|u(t)|^{2} d x-\frac{6 p}{p+2}\|\psi(t)\|_{L^{p+2}}^{p+2}  \tag{4.10}\\
& \quad+C C_{1}^{\frac{3 p}{2 s}}\|\psi(t)\|_{L^{2}(|x| \geq R)}^{\frac{4 s-(3-2 s) p}{2 s}}+C C_{1}^{\frac{3-2 r}{s}}\|\psi(t)\|_{L^{2}(|x| \geq R)}^{\frac{4 s+2 r-3}{2 s}} \\
& \leq 4 Q(\psi(t))+C R^{-2 s}\|\psi(t)\|_{L^{2}(|x| \geq R)}^{2}+C C_{1}^{\frac{3 p}{2 s}}\|\psi(t)\|_{L^{2}(|x| \geq R)}^{\frac{4 s-(3-2 s) p}{2 s}} \\
& \quad+C C_{1}^{\frac{3-2 r}{s}}\|\psi(t)\|_{L^{2}(|x| \geq R)}^{\frac{4 s+2 r-3}{2 s}} .
\end{align*}
$$

By Lemma 4.1. we see that for any $\eta>0$ and any $R>1$, there exists $C>0$ independent of $R$ and $C_{1}$ such that for any $t \in\left[0, T_{0}\right]$ with $T_{0}=\frac{\eta R}{C C_{1}^{2}}$, we have

$$
\begin{aligned}
\frac{d}{d t} \mathcal{M}_{\varphi_{R}}[\psi(t)] \leq & 4 Q(\psi(t))+C R^{-2 s}\left(\eta+o_{R}(1)\right)^{2}+C C_{1}^{\frac{3 p}{2 s}}\left(\eta+o_{R}(1)\right)^{\frac{4 s-p(3-2 s)}{2 s}} \\
& +C C_{1}^{\frac{3-2 r}{s}}\left(\eta+o_{R}(1)\right)^{\frac{4 s+2 r-3}{2 s}} \\
\leq & -4 \delta+C R^{-2 s}\left(\eta^{2}+o_{R}(1)\right)+C C_{1}^{\frac{3 p}{2 s}}\left(\eta^{\frac{4 s-p(3-2 s)}{2 s}}+o_{R}(1)\right) \\
+ & C C_{1}^{\frac{3-2 r}{s}}\left(\eta^{\frac{4 s+2 r-3}{2 s}}+o_{R}(1)\right) .
\end{aligned}
$$

We first choose $\eta>0$ small enough so that

$$
C C_{1}^{\frac{3 p}{2 s}} \eta^{\frac{4 s-p(3-2 s)}{2 s}}+C C_{1}^{\frac{3-2 r}{s}} \eta^{\frac{4 s+2 r-3}{2 s}}=-3 \delta>0
$$

We next choose $R>1$ large enough so that

$$
\begin{equation*}
\frac{d}{d t} \mathcal{M}_{\varphi_{R}}[\psi(t)] \leq-\delta<0 \tag{4.11}
\end{equation*}
$$

for any $t \in\left[0, T_{0}\right]$ with $T_{0}=\frac{\eta R}{C C_{1}^{2}}$. Note that $\eta>0$ is fixed, so we can choose $R>1$ large enough such that $T_{0}$ is as large as we want. From 4.11) it follows that

$$
\mathcal{M}_{\varphi_{R}}[\psi(t)] \leq-\delta t
$$

for all $t \in\left[t_{0}, T_{0}\right]$ with some sufficiently large $t_{0} \in\left[0, T_{0}\right]$. On the other hand, by Lemma 3.1 and the conservation of mass, we have for any $t \in[0,+\infty)$,

$$
\begin{aligned}
\left|\mathcal{M}_{\varphi_{R}}[\psi(t)]\right| & \leq C C\left(\varphi_{R}\right)\left(\left\||\nabla|^{1 / 2} \psi(t)\right\|_{L^{2}}^{2}+\|\psi(t)\|_{L^{2}}\left\||\nabla|^{1 / 2} \psi(t)\right\|_{L^{2}}\right) \\
& \leq C C\left(\varphi_{R}\right)\left(\left\||\nabla|^{1 / 2} \psi(t)\right\|_{L^{2}}^{2}+\|\psi(t)\|_{L^{2}}^{2}\right) \\
& \leq C C\left(\varphi_{R}\right)\left(\left\||\nabla|^{1 / 2} \psi(t)\right\|_{L^{2}}^{2}+1\right)
\end{aligned}
$$

By interpolating between $L^{2}$ and $\dot{H}^{s}$, we obtain for any $t \in\left[t_{0}, T_{0}\right]$

$$
\delta t \leq-\mathcal{M}_{\varphi_{R}}[\psi(t)]=\left|\mathcal{M}_{\varphi_{R}}[\psi(t)]\right| \leq C C\left(\varphi_{R}\right)\left(\left\|(-\Delta)^{s / 2} \psi(t)\right\|_{L^{2}}^{\frac{1}{s}}+1\right)
$$

This implies that

$$
\begin{equation*}
\left\|(-\Delta)^{s / 2} \psi(t)\right\|_{L^{2}} \geq C t^{s} \tag{4.12}
\end{equation*}
$$

for all $t \in\left[t_{1}, T_{0}\right]$ with some sufficiently large $t_{1} \in\left[t_{0}, T_{0}\right]$. Taking $t$ close to $T_{0}=\frac{\eta R}{C C_{1}^{2}}$, we see that $\left\|(-\Delta)^{s / 2} \psi(t)\right\|_{L^{2}} \rightarrow \infty$ as $R \rightarrow \infty$. Taking $R>1$ sufficiently large, we have a contradiction with 4.1. The proof is complete.

## 5. Strong instability of standing waves

In this section, we prove Theorem 1.2 Let us start with the following characterization of the ground state related to (1.12).

Proposition 5.1. Let $\omega>0,2 s+2 r>3$ and $\frac{4 s}{3} \leq p<\frac{4 s}{3-2 s}$. Then $u$ is the ground state related to 1.12 if and only if $u$ solves the minimization problem

$$
\begin{equation*}
d(\omega)=\inf \left\{S_{\omega}(v): v \in H^{s} \backslash\{0\}, K_{\omega}(v)=0\right\} \tag{5.1}
\end{equation*}
$$

To solve this minimization problem, we consider the minimization problem

$$
\begin{equation*}
\widetilde{d}(\omega)=\inf \left\{\widetilde{S}_{\omega}(v): v \in H^{s} \backslash\{0\}, K_{\omega}(v) \leq 0\right\} \tag{5.2}
\end{equation*}
$$

where

$$
\begin{align*}
\widetilde{S}_{\omega}(v) & :=S_{\omega}(v)-\frac{K_{\omega}(v)}{4 s+2 r-3} \\
& =\frac{\omega s}{4 s+2 r-3}\|v\|_{L^{2}}^{2}+\frac{p(s+r)-2 s}{(p+2)(4 s+2 r-3)}\|v\|_{L^{p+2}}^{p+2} \tag{5.3}
\end{align*}
$$

If $K_{\omega}(v)<0$, then

$$
\begin{aligned}
K_{\omega}(\lambda v)= & \frac{4 s+2 r-3}{2} \lambda^{2}\|v\|_{\dot{H}^{s}}^{2}+\frac{4 s+2 r-3}{4} \lambda^{4} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x \\
& +\frac{\omega(2 s+2 r-3)}{2} \lambda^{2}\|v\|_{L^{2}}^{2}-\frac{(s+r)(p+2)-3}{p+2} \lambda^{p+2}\|v\|_{L^{p+2}}^{p+2}>0
\end{aligned}
$$

for sufficiently small $\lambda>0$. Thus, there exists $\lambda_{0} \in(0,1)$ such that $K_{\omega}\left(\lambda_{0} v\right)=0$. It follows that

$$
\widetilde{S}_{\omega}\left(\lambda_{0} v\right)=\frac{\omega s}{4 s+2 r-3} \lambda_{0}^{2}\|v\|_{L^{2}}^{2}+\frac{p(s+r)-2 s}{(p+2)(4 s+2 r-3)} \lambda_{0}^{p+2}\|v\|_{L^{p+2}}^{p+2}<\widetilde{S}_{\omega}(v)
$$

This implies that

$$
\begin{equation*}
\widetilde{d}(\omega)=\inf \left\{\widetilde{S}_{\omega}(v): v \in H^{s} \backslash\{0\}, K_{\omega}(v)=0\right\} \tag{5.4}
\end{equation*}
$$

In following lemma, we will solve the minimizing problem 5.2).
Lemma 5.2. Let $\omega>0,2 s+2 r>3$ and $\frac{4 s}{3} \leq p<\frac{4 s}{3-2 s}$. Then there exists $u \in H^{s} \backslash\{0\}$, such that $K_{\omega}(u)=0$ and $\widetilde{S}_{\omega}(u)=\widetilde{d}(\omega)$.

Proof. We first show that $\tilde{d}(\omega)>0$. From $K_{\omega}(v) \leq 0$, we have

$$
\begin{aligned}
& \frac{4 s+2 r-3}{2}\|v\|_{\dot{H}^{s}}^{2}+\frac{\omega(2 s+2 r-3)}{2}\|v\|_{L^{2}}^{2}+\frac{4 s+2 r-3}{4} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x \\
& \leq \frac{(s+r)(p+2)-3}{p+2}\|v\|_{L^{p+2}}^{p+2}
\end{aligned}
$$

which implies

$$
\frac{1}{2} H_{\omega}(v) \leq \frac{(s+r)(p+2)-3}{(p+2)(2 s+2 r-3)} H_{\omega}(v)^{\frac{p}{2}+1}
$$

where $H_{\omega}(v)=\|v\|_{\dot{H}^{s}}^{2}+\omega\|v\|_{L^{2}}^{2}$. Thus, there exists $C_{0}>0$ such that $H_{\omega}(v)>C_{0}$ for all $K_{\omega}(v) \leq 0$. This implies that there exists $C_{1}>0$ such that

$$
\begin{aligned}
\widetilde{S}_{\omega}(v) & \geq \frac{p(s+r)-2 s}{2(4 s+2 r-3)}\|v\|_{L^{p+2}}^{p+2} \\
& \geq \frac{p(s+r)-2 s}{(p+2)(4 s+2 r-3)} \frac{(2 s+2 r-3)}{(s+r)(p+2)-3} H_{\omega}(v) \geq C_{1}
\end{aligned}
$$

Taking the infimum over $v$, we obtain $\widetilde{d}(\omega)>0$.
We now show that the minimizing problem (5.2) attains its minimum. Let $\left\{v_{n}\right\}$ be a minimizing sequence for 55.2 , i.e., $\left\{v_{n}\right\} \subseteq H^{s} \backslash\{0\}, K_{\omega}\left(v_{n}\right) \leq 0$ and $\widetilde{S}_{\omega}\left(v_{n}\right) \rightarrow \widetilde{d}(\omega)$ as $n \rightarrow \infty$. Thus, there exists $C>0$ such that

$$
\begin{equation*}
\left\|v_{n}\right\|_{L^{2}}^{2}+\left\|v_{n}\right\|_{L^{p+2}}^{p+2} \leq C \tag{5.5}
\end{equation*}
$$

This, together with $K_{\omega}\left(v_{n}\right) \leq 0$ implies that $\left\{v_{n}\right\}$ is bounded in $H^{s}$. It follows from $\tilde{d}(\omega)>0$ that $\liminf _{n \rightarrow \infty}\left\|v_{n}\right\|_{L^{p+2}}^{p+2}>0$. Therefore, applying Lemma 2.4 , there exists a subsequence, still denoted by $\left\{v_{n}\right\}$ and $u \in H^{s} \backslash\{0\}$ such that

$$
u_{n}:=\tau_{x_{n}} v_{n} \rightharpoonup u \neq 0 \text { weakly in } H^{s}
$$

for some $\left\{x_{n}\right\} \subseteq \mathbb{R}^{3}$. We deduce from Brezis-Lieb's lemma (Lemma 2.2) and Lemma 2.3 that

$$
\begin{gather*}
K_{\omega}\left(u_{n}\right)-K_{\omega}\left(u_{n}-u\right)-K_{\omega}(u) \rightarrow 0  \tag{5.6}\\
\widetilde{S}_{\omega}\left(u_{n}\right)-\widetilde{S}_{\omega}\left(u_{n}-u\right)-\widetilde{S}_{\omega}(u) \rightarrow 0 \tag{5.7}
\end{gather*}
$$

Now, we claim that $K_{\omega}(u) \leq 0$. If not, it follows from 5.6) and $K_{\omega}\left(u_{n}\right) \leq 0$ that $K_{\omega}\left(u_{n}-u\right) \leq 0$ for sufficiently large $n$. Thus, by the definition of $\widetilde{d}(\omega)$, it follows that

$$
\widetilde{S}_{\omega}\left(u_{n}-u\right) \geq \widetilde{d}(\omega)
$$

which, together with $\widetilde{S}_{\omega}\left(u_{n}\right) \rightarrow \widetilde{d}(\omega)$, implies that $\widetilde{S}_{\omega}(u) \leq 0$, which is a contradiction with $\widetilde{S}_{\omega}(u)>0$. We thus obtain $K_{\omega}(u) \leq 0$.

Furthermore, we deduce from the definition of $\widetilde{d}(\omega)$ and the weak lower semicontinuity of norm that

$$
\widetilde{d}(\omega) \leq \widetilde{S}_{\omega}(u) \leq \liminf _{n \rightarrow \infty} \widetilde{S}_{\omega}\left(u_{n}\right)=\widetilde{d}(\omega)
$$

This yields $\widetilde{S}_{\omega}(u)=\widetilde{d}(\omega)$.
Finally, we show that $K_{\omega}(u)=0$. Suppose that $K_{\omega}(u)<0$ and set

$$
\begin{aligned}
K_{\omega}\left(u^{\lambda}\right)= & \frac{4 s+2 r-3}{2} \lambda^{2 s}\|u\|_{\dot{H}^{s}}^{2}+\frac{\omega(2 s+2 r-3)}{2}\|u\|_{L^{2}}^{2} \\
& +\frac{4 s+2 r-3}{4} \lambda^{3-2 r} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|u|^{2}\right)|u|^{2} d x \\
& -\frac{(s+r)(p+2)-3}{p+2} \lambda^{\frac{3 p}{2}}\|u\|_{L^{p+2}}^{p+2}>0
\end{aligned}
$$

for sufficiently small $\lambda>0$. Then there exists $\lambda_{0} \in(0,1)$ such that $K_{\omega}\left(u^{\lambda_{0}}\right)=0$. It follows that

$$
\begin{aligned}
\widetilde{S}_{\omega}\left(u^{\lambda_{0}}\right) & =\frac{\omega s}{4 s+2 r-3}\|u\|_{L^{2}}^{2}+\frac{p(s+r)-2 s}{(p+2)(4 s+2 r-3)} \lambda_{0}^{\frac{3 p}{2}}\|u\|_{L^{p+2}}^{p+2} \\
& <\widetilde{S}_{\omega}(u)=\widetilde{d}(\omega)
\end{aligned}
$$

which contradicts the definition of $\widetilde{d}(\omega)$. Hence, we have $K_{\omega}(u)=0$.
From $d(\omega)=\widetilde{d}(\omega)$ and the above lemma, we can obtain the existence of minimization problem (5.1).

Lemma 5.3. Let $\omega>0,2 s+2 r>3$ and $\frac{4 s}{3} \leq p<\frac{4 s}{3-2 s}$. Then there exists $u \in H^{s} \backslash\{0\}$ such that $K_{\omega}(u)=0$ and $S_{\omega}(u)=d(\omega)$.
Lemma 5.4. Let $\omega>0,2 s+2 r>3$, and $\frac{4 s}{3} \leq p<\frac{4 s}{3-2 s}$. Assume that $u \in$ $H^{s} \backslash\{0\}$ is a solution of the minimizing problem 5.1), i.e., such that $K_{\omega}(u)=0$ and $S_{\omega}(u)=d(\omega)$. Then $S_{\omega}^{\prime}(u)=0$.
Proof. We firstly prove $K_{\omega}^{\prime}(u) \neq 0$. If $K_{\omega}^{\prime}(u)=0$, then we have

$$
\begin{align*}
& (4 s+2 r-3)(-\Delta)^{s} u+\omega(2 s+2 r-3) u+(4 s+2 r-3)\left(|x|^{-(3-2 r)} *|u|^{2}\right) u \\
& -((s+r)(p+2)-3)|u|^{p} u=0 \tag{5.8}
\end{align*}
$$

Then

$$
\begin{gather*}
A+B+C-D=d(\omega), \\
(4 s+2 r-3) A+(2 s+2 r-3) B+(4 s+2 r-3) C \\
-((s+r)(p+2)-3) D=0 \\
2(4 s+2 r-3) A+2(2 s+2 r-3) B+4(4 s+2 r-3) C  \tag{5.9}\\
-(p+2)((s+r)(p+2)-3) D=0, \\
(3-2 s)(4 s+2 r-3) A+3(2 s+2 r-3) B+(4 s+2 r-3)(3+2 r) C \\
-3((s+r)(p+2)-3) D=0,
\end{gather*}
$$

where

$$
\begin{gathered}
A=\frac{1}{2}\|u\|_{\dot{H}^{s}}^{2}, \quad B=\frac{\omega}{2}\|u\|_{L^{2}}^{2} \\
C=\frac{1}{4} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|u|^{2}\right)(x)|u(x)|^{2} d x, \quad D=\frac{1}{p+2}\|u\|_{L^{p+2}}^{p+2} .
\end{gathered}
$$

The first equation comes from the fact that $S_{\omega}(u)=d(\omega)$. The second one holds since $K_{\omega}(u)=0$. The third one follows by multiplying (5.8) by $u$ and integrating both sides. The fourth one is derived by applying the Pohozaev equality to (5.8).

After a direct calculations, we have

$$
\begin{gathered}
s A=t C, \quad C=\frac{p((s+r)(p+2)-3) D}{2(4 s+2 r-3)} \\
(2 s+2 r-3) B+\frac{(p(s+r)-2 s)((s+r)(p+2)-3)}{2 s} D=0
\end{gathered}
$$

These $A=B=C=D=0$ which is a contradiction with $A, B, C, D>0$. Thus, $K_{\omega}^{\prime}(u) \neq 0$.

Next, applying the Lagrange multiplier rule, there exists $\mu \in \mathbb{R}$ such that $S_{\omega}^{\prime}(u)+$ $\mu K_{\omega}^{\prime}(u)=0$. We claim that $\mu=0$. As above, the equation $S_{\omega}^{\prime}(u)+\mu K_{\omega}^{\prime}(u)=0$ can be written as

$$
\begin{align*}
& (-\Delta)^{s} u+\omega u+\left(|x|^{-(3-2 r)} *|u|^{2}\right) u-|u|^{p} u \\
& +\mu\left[(4 s+2 r-3)\left(|x|^{-(3-2 r)} *|u|^{2}\right) u+\omega(2 s+2 r-3) u\right.  \tag{5.10}\\
& \left.+(4 s+2 r-3)(-\Delta)^{s} u-((s+r)(p+2)-3)|u|^{p} u\right]=0
\end{align*}
$$

By the same argument as in (5.9), we have

$$
\begin{gathered}
A+B+C-D=d(\omega) \\
(4 s+2 r-3) A+(2 s+2 r-3) B+(4 s+2 r-3) C-((s+r)(p+2)-3) D=0 \\
2(\mu(4 s+2 r-3)+1) A+2(\mu(2 s+2 r-3)+1) B \\
+4(\mu(4 s+2 r-3)+1) C-(p+2)(\mu((s+r)(p+2)-3)+1) D=0 \\
(3-2 s)(\mu(4 s+2 r-3)+1) A+3(\mu(2 s+2 r-3)+1) B \\
+(\mu(4 s+2 r-3)+1)(3+2 r) C-3(\mu((s+r)(p+2)-3)+1) D=0
\end{gathered}
$$

We now deal with the above system. Consider $A, B, C, D$ as unknown quantities, and denote the coefficient matrix by $M$. Computing its determinant, we have

$$
\operatorname{det} M=-4 s \mu p(s+r)(1+\mu(4 s+2 r-3))((p-2) s+p r)
$$

Note that

$$
\operatorname{det} M=0 \Longleftrightarrow \mu=0, p=0, \mu=-\frac{1}{4 s+2 r-3},(p-2) s+p r=0
$$

Because of $2 s+2 r>3$ and $\frac{4 s}{3} \leq p<\frac{4 s}{3-2 s}$, it follows that $(p-2) s+p r>0$. We will show that $\mu$ must be equal to zero by excluding the other possibilities:
(1) If $\mu \neq 0, \mu \neq-\frac{1}{4 s+2 r-3}$, then $\operatorname{det} M \neq 0$, and hence the linear system has a unique solution (depending on the parameters $\mu, p, d(\omega)$ ). Applying Cramer's rule, we obtain

$$
D=-\frac{d(\omega)(4 s+2 r-3)(2 s+2 r-3)}{p(s+r)((p-2) s+p r)}<0
$$

which contradicts $D>0$.
(2) If $\mu=-\frac{1}{4 s+2 r-3}$, then the third equation reads

$$
4 s B+(p+2)(s(p-2)+p r) D=0
$$

which contradicts $B, D>0$. Thus, $\mu=0$ and $S_{\omega}^{\prime}(u)=0$.

We now denote the set of all minimizers of 5.1 by

$$
\mathcal{M}_{\omega}=\left\{u \in H^{s} \backslash\{0\}: S_{\omega}(u)=d(\omega), K_{\omega}(u)=0\right\}
$$

Lemma 5.5. $\mathcal{M}_{\omega} \subseteq \mathcal{G}_{\omega}$.
Proof. Let $u \in \mathcal{M}_{\omega}$. It follows from Lemma 5.4 that $S_{\omega}^{\prime}(u)=0$. In particular, we have $u \in \mathcal{A}_{\omega}$. To prove $u \in \mathcal{G}_{\omega}$, it remains to show that $S_{\omega}(u) \leq S_{\omega}(v)$ for all $v \in \mathcal{A}_{\omega}$. To see this, we notice that

$$
K_{\omega}(v)=(s+r)\left\langle S_{\omega}^{\prime}(v), v\right\rangle-I_{\omega}(v)=0
$$

for all $v \in \mathcal{A}_{\omega}$, where $I_{\omega}(v)$ is defined by 2.3. By definition of $d(\omega)$, we have $S_{\omega}(u) \leq S_{\omega}(v)$. Thus, $u \in \mathcal{G}_{\omega}$.

Lemma 5.6. $\mathcal{G}_{\omega} \subset \mathcal{M}_{\omega}$.
Proof. Let $u \in \mathcal{G}_{\omega}$. Since $\mathcal{M}_{\omega}$ is not empty, we take $v \in \mathcal{M}_{\omega}$. By Lemma 5.5 , $v \in \mathcal{G}_{\omega}$. In particular, $S_{\omega}(u)=S_{\omega}(v)$. Since $v \in \mathcal{M}_{\omega}$, we obtain

$$
S_{\omega}(u)=S_{\omega}(v)=d(\omega)
$$

It remains to show that $K_{\omega}(u)=0$. Since $u \in \mathcal{A}_{\omega}$, we have $S_{\omega}^{\prime}(u)=0$ and $I_{\omega}(u)=0$, hence $K_{\omega}(u)=(s+r)\left\langle S_{\omega}^{\prime}(u), u\right\rangle-I_{\omega}(u)=0$. Therefore, $u \in \mathcal{M}_{\omega}$.

Proof of Proposition 5.1. It follows immediately from Lemmas 5.3, 5.5, and 5.6.
When $\frac{4 s}{3} \leq p<\frac{4 s}{3-2 s}$, to study the strong instability of standing waves for (1.4), we need to establish the following characterization of the ground state related to (1.12).

Lemma 5.7. Let $\omega>0,2 s+2 r>3, \frac{4 s}{3} \leq p<\frac{4 s}{3-2 s}$, and $u$ be the ground state related to 1.12. Then

$$
\begin{equation*}
S_{\omega}(u)=\inf \left\{S_{\omega}(v): v \in H^{s} \backslash\{0\}, Q(v)=0\right\} \tag{5.11}
\end{equation*}
$$

Proof. Firstly, we claim that the minimizing problem in 5.11 is well-defined. Let $v \in H^{s} \backslash\{0\}$ and $Q(v)=0$. If $p=\frac{4 s}{3}$, then

$$
\begin{align*}
S_{\omega}(v) & =S_{\omega}(v)-\frac{1}{2 s} Q(v) \\
& =\frac{\omega}{2}\|v\|_{L^{2}}^{2}+\frac{2 s+2 r-3}{8 s} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x>0 . \tag{5.12}
\end{align*}
$$

And if $\frac{4 s}{3}<p<\frac{4 s}{3-2 s}$, then

$$
\begin{align*}
& S_{\omega}(v) \\
& =S_{\omega}(v)-\frac{2}{3 p} Q(v)  \tag{5.13}\\
& =\frac{3 p-4 s}{6 p}\|v\|_{\dot{H}^{s}}^{2}+\frac{\omega}{2}\|v\|_{L^{2}}^{2}+\frac{3 p+4 t-6}{12 p} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x>0
\end{align*}
$$

Thus we denote $d:=\inf \left\{S_{\omega}(v): v \in H^{s} \backslash\{0\}, Q(v)=0\right\}$. Firstly, we deduce from (2.2) and (2.3) that

$$
K_{\omega}(u)=Q(u)=0
$$

By the definition of $d$, we have

$$
S_{\omega}(u) \geq d
$$

Let $v \in H^{s} \backslash\{0\}$ be such that $Q(v)=0$. If $K_{\omega}(v)=0$, then it follows from Proposition 5.1 that

$$
S_{\omega}(v) \geq S_{\omega}(u)
$$

If $K_{\omega}(v) \neq 0$, we notice that

$$
\begin{aligned}
K_{\omega}\left(v^{\lambda}\right)= & \frac{4 s+2 r-3}{2} \lambda^{2 s}\|v\|_{\dot{H}^{s}}^{2}+\frac{\omega(2 s+2 r-3)}{2}\|v\|_{L^{2}}^{2} \\
& +\frac{4 s+2 r-3}{4} \lambda^{3-2 r} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x \\
& -\frac{(s+r)(p+2)-3}{p+2} \lambda^{\frac{3 p}{2}}\|v\|_{L^{p+2}}^{p+2},
\end{aligned}
$$

where $v^{\lambda}(x):=\lambda^{3 / 2} v(\lambda x)$. When $\frac{4 s}{3}<p<\frac{4 s}{3-2 s}$, we have

$$
\begin{equation*}
\lim _{\lambda \rightarrow 0} K_{\omega}\left(v^{\lambda}\right)=\frac{\omega(2 s+2 r-3)}{2}\|v\|_{L^{2}}^{2}>0, \quad \text { and } \quad \lim _{\lambda \rightarrow \infty} K_{\omega}\left(v^{\lambda}\right)<0 \tag{5.14}
\end{equation*}
$$

When $p=4 s / 3$, it follows from $Q(v)=0$ that

$$
s\|v\|_{\dot{H}^{s}}^{2}<\frac{3 p}{2(p+2)}\|v\|_{L^{p+2}}^{p+2}
$$

which implies that 5.14 holds. Thus, there exists $\lambda_{0}>0$ such that $K_{\omega}\left(v^{\lambda_{0}}\right)=0$. This implies that

$$
S_{\omega}\left(v^{\lambda_{0}}\right) \geq S_{\omega}(u)
$$

On the other hand, by some basic calculations, we have

$$
\begin{aligned}
\partial_{\lambda} S_{\omega}\left(v^{\lambda}\right)= & s \lambda^{2 s-1}\|v\|_{\dot{H}^{s}}^{2}+\frac{3-2 r}{4} \lambda^{2-2 r} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x \\
& -\frac{\lambda^{\frac{3 p}{2}-1}}{p+2} \frac{3 p}{2}\|v\|_{L^{p+2}}^{p+2} \\
= & \frac{Q\left(v^{\lambda}\right)}{\lambda}
\end{aligned}
$$

Next, we define

$$
\begin{aligned}
f(\lambda) & :=Q\left(v^{\lambda}\right) \\
& =s \lambda^{2 s}\|v\|_{\dot{H}^{s}}^{2}+\frac{3-2 r}{4} \lambda^{3-2 r} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x-\frac{\lambda^{\frac{3 p}{2}}}{p+2} \frac{3 p}{2}\|v\|_{L^{p+2}}^{p+2} .
\end{aligned}
$$

When $p=\frac{4 s}{3}$, it follows from $Q(v)=0$ that $s\|v\|_{\dot{H}^{s}}^{2}<\frac{3 p}{2(p+2)}\|v\|_{L^{p+2}}^{p+2}$. Thus, it is easy to see that the equation $f(\lambda)=0$ admits a unique positive solution $\lambda=1$.

When $\frac{4 s}{3}<p<\frac{4 s}{3-2 s}$, assume that there exists $\lambda_{1} \neq 1$ such that $f\left(\lambda_{1}\right)=0$. It easily follows that

$$
\frac{3-2 r}{4} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x\left(\lambda_{1}^{2 s}-\lambda_{1}^{3-2 r}\right)=\frac{\|v\|_{L^{p+2}}^{p+2}}{p+2} \frac{3 p}{2}\left(\lambda_{1}^{2 s}-\lambda_{1}^{\frac{3 p}{2}}\right)
$$

If $\lambda_{1}>1$, then $\lambda_{1}^{2 s}-\lambda_{1}^{3-2 r}>0$ and $\lambda_{1}^{2 s}-\lambda_{1}^{\frac{3 p}{2}}<0$, which is a contradiction. If $\lambda_{1}<1$, then $\lambda_{1}^{2 s}-\lambda_{1}^{3-2 r}<0$ and $\lambda_{1}^{2 s}-\lambda_{1}^{\frac{3 p}{2}}>0$, which is a contradiction. Therefore, the equation $f(\lambda)=0$ admits a unique positive solution $\lambda=1$. Therefore,

$$
\begin{gathered}
\partial_{\lambda} S_{\omega}\left(v^{\lambda}\right)>0, \quad \text { for all } \lambda \in(0,1) \\
\partial_{\lambda} S_{\omega}\left(v^{\lambda}\right)<0, \quad \text { for all } \lambda \in(1, \infty)
\end{gathered}
$$

We thus obtain that $S_{\omega}\left(v^{\lambda}\right)<S_{\omega}(v)$ for any $\lambda>0$ and $\lambda \neq 1$. In particular, we have $S_{\omega}\left(v^{\lambda_{0}}\right) \leq S_{\omega}(v)$. Thus, $S_{\omega}(u) \leq S_{\omega}\left(v^{\lambda_{0}}\right) \leq S_{\omega}(v)$ for all $v \in H^{s} \backslash\{0\}$ and $Q(v)=0$. Taking the infimum over $v$, we have $S_{\omega}(u) \leq d$. This completes the proof.

To obtain the key estimate (5.19), we need establish the following variational characterization of the ground states to 1.12 . Firstly, when $p=\frac{4 s}{3}$, we define

$$
\begin{align*}
S_{\omega}^{1}(v) & :=S_{\omega}(v)-\frac{1}{2 s} Q(v) \\
& =\frac{\omega}{2}\|v\|_{L^{2}}^{2}+\frac{2 s+2 r-3}{8 s} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x \tag{5.15}
\end{align*}
$$

When $\frac{4 s}{3}<p<\frac{4 s}{3-2 s}$, we define

$$
\begin{align*}
S_{\omega}^{2}(v):= & S_{\omega}(v)-\frac{2}{3 p} Q(v) \\
= & \frac{3 p-4 s}{6 p}\|v\|_{\dot{H}^{s}}^{2}+\frac{\omega}{2}\|v\|_{L^{2}}^{2}  \tag{5.16}\\
& +\frac{3 p+4 t-6}{12 p} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x .
\end{align*}
$$

Lemma 5.8. Let $\omega>0,2 s+2 r>3, \frac{4 s}{3} \leq p<\frac{4 s}{3-2 s}$, and $u$ be the ground state related to 1.12 . Then for $k=1,2$ we have

$$
\begin{equation*}
S_{\omega}(u)=S_{\omega}^{k}(u)=\inf \left\{S_{\omega}^{k}(v): v \in H^{s} \backslash\{0\}, Q(v) \leq 0\right\} \tag{5.17}
\end{equation*}
$$

Proof. We only prove the case $k=1$. The proof of the case $k=2$ is similar. We denote

$$
d^{1}(\omega)=\inf \left\{S_{\omega}^{k}(v): v \in H^{s} \backslash\{0\}, Q(v) \leq 0\right\}
$$

Since $u$ is the ground state related to $1.12, Q(u)=0$. It follows from the definition of $d^{1}(\omega)$ that

$$
\begin{equation*}
S_{\omega}^{1}(u) \geq d^{1}(\omega) \tag{5.18}
\end{equation*}
$$

Let $v \in H^{s} \backslash\{0\}$ and $Q(v) \leq 0$. If $Q(v)=0$, then from Lemma 5.7 it follows that

$$
S_{\omega}^{1}(v)=S_{\omega}(v)-\frac{1}{2 s} Q(v)=S_{\omega}(v) \geq S_{\omega}(u)=S_{\omega}^{1}(u)
$$

If $Q(v)<0$, we note that

$$
Q\left(v^{\lambda}\right)=\lambda^{2 s}\|v\|_{\dot{H}^{s}}^{2}+\frac{\lambda^{3-2 r}}{4} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x-\frac{\lambda^{\frac{3 p}{2}}}{p+2} \frac{3 p}{2}\|v\|_{L^{p+2}}^{p+2}>0
$$

for sufficiently small $\lambda>0$, so there exists $\lambda_{0} \in(0,1)$ such that $Q\left(v^{\lambda_{0}}\right)=0$. We thus have

$$
\begin{aligned}
S_{\omega}^{1}(v) & =\frac{\omega}{2}\|v\|_{L^{2}}^{2}+\frac{2 s+2 r-3}{8 s} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x \\
& \geq \frac{\omega}{2}\|v\|_{L^{2}}^{2}+\frac{2 s+2 r-3}{8 s} \lambda_{0}^{3-2 r} \int_{\mathbb{R}^{3}}\left(|x|^{-(3-2 r)} *|v|^{2}\right)|v|^{2} d x \\
& =S_{\omega}^{1}\left(v^{\lambda_{0}}\right)=S_{\omega}\left(v^{\lambda_{0}}\right) \geq S_{\omega}(u)=S_{\omega}^{1}(u)
\end{aligned}
$$

This implies that $d^{1}(\omega) \geq S_{\omega}^{1}(u)$. This, together with 5.18 implies that $S_{\omega}^{1}(u)=$ $d^{1}(\omega)$.

Let $u$ be the ground state related to 1.12 . We define

$$
\mathcal{B}_{\omega}=\left\{v \in H^{s} \backslash\{0\}: S_{\omega}(v)<S_{\omega}(u), Q(v)<0\right\} .
$$

Lemma 5.9. Let $\omega>0,2 s+2 r>3$, and $u$ be the ground state related to 1.12. If $\frac{4 s}{3} \leq p<\frac{4 s}{3-2 s}$, then the set $\mathcal{B}_{\omega}$ is invariant under the flow of 1.4. That is, if $\psi_{0} \in \mathcal{B}_{\omega}$, then the solution $\psi(t)$ to (1.4) with initial data $\psi_{0}$ belongs to $\mathcal{B}_{\omega}$ and

$$
\begin{equation*}
Q(\psi(t)) \leq 2 s\left(S\left(\psi_{0}\right)-S(u)\right) \tag{5.19}
\end{equation*}
$$

for any $t \in\left[0, T^{*}\right)$.
Proof. Let $\psi_{0} \in \mathcal{B}_{\omega}$, by Proposition 2.1, we see that there exists a unique solution $\psi \in C\left(\left[0, T^{*}\right), H^{s}\right)$ with initial data $\psi_{0}$. We deduce from the conservations of mass and energy that

$$
\begin{equation*}
S_{\omega}(\psi(t))=S_{\omega}\left(\psi_{0}\right)<S_{\omega}(u) \tag{5.20}
\end{equation*}
$$

for any $t \in\left[0, T^{*}\right)$. In addition, by the continuity of the function $t \mapsto Q(\psi(t))$ and Lemma 5.7, if there exists $t_{0} \in\left[0, T^{*}\right)$ such that $Q\left(\psi\left(t_{0}\right)\right)=0$, then $S_{\omega}\left(\psi\left(t_{0}\right)\right) \geq$ $S_{\omega}(u)$, which contradicts 5.20 . Therefore, we have $Q(\psi(t))<0$ for any $t \in\left[0, T^{*}\right)$. This, together with Lemma 5.8 implies that

$$
\begin{aligned}
& S_{\omega}(u) \leq S_{\omega}^{1}(\psi(t))=S_{\omega}(\psi(t))-\frac{1}{2 s} Q(\psi(t))=S_{\omega}\left(\psi_{0}\right)-\frac{Q(\psi(t))}{2 s} \\
& S_{\omega}(u) \leq S_{\omega}^{2}(\psi(t))=S_{\omega}(\psi(t))-\frac{2}{3 p} Q(\psi(t))<S_{\omega}\left(\psi_{0}\right)-\frac{Q(\psi(t))}{2 s}
\end{aligned}
$$

for all $t \in\left[0, T^{*}\right)$. This completes the proof.
Proof of Theorem 1.2. Let $u$ be the ground state related to 1.12 and $\left\{\lambda_{n}\right\} \subseteq \mathbb{R}^{+}$ be such that $\lambda_{n}>1$ and $\lim _{n \rightarrow \infty} \lambda_{n}=1$. We take the initial data

$$
\psi_{0, n}(x):=\lambda_{n}^{3 / 2} u\left(\lambda_{n} x\right)
$$

Therefore,

$$
\begin{gathered}
\lim _{n \rightarrow \infty}\left\|\psi_{0, n}\right\|_{L^{2}}=\lim _{n \rightarrow \infty}\|u\|_{L^{2}}=\|u\|_{L^{2}} \\
\lim _{n \rightarrow \infty}\left\|\psi_{0, n}\right\|_{\dot{H}^{s}}=\lim _{n \rightarrow \infty} \lambda_{n}^{s}\|u\|_{\dot{H}^{s}}=\|u\|_{\dot{H}^{s}}
\end{gathered}
$$

Thus, we deduce from Brezis-Lieb's lemma (Lemma 2.2) that $\psi_{0, n} \rightarrow u$ in $H^{s}$ as $n \rightarrow \infty$. By Lemma 5.7, we have

$$
S_{\omega}\left(\psi_{0, n}\right)<S_{\omega}(u), Q\left(\psi_{0, n}\right)<0
$$

for all $n \geq 1$. Thus, $\psi_{0, n} \in \mathcal{B}_{\omega}$. Let $\psi_{n}$ be the maximal solution of 1.4 with the initial data $\psi_{0, n}$. We deduce from Lemma 5.9 that $\psi_{n}(t) \in \mathcal{B}_{\omega}$ for all $t \in\left[0, T^{*}\right)$ and

$$
Q\left(\psi_{n}(t)\right) \leq 2 s\left(S\left(\psi_{0, n}\right)-S(u)\right)<0
$$

Thus, applying Theorem 1.1, we obtain that the solution $\psi_{n}(t)$ of (1.4) with initial data $\psi_{0, n}$ blows up in finite or infinite time for any $n \geq 1$.

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